



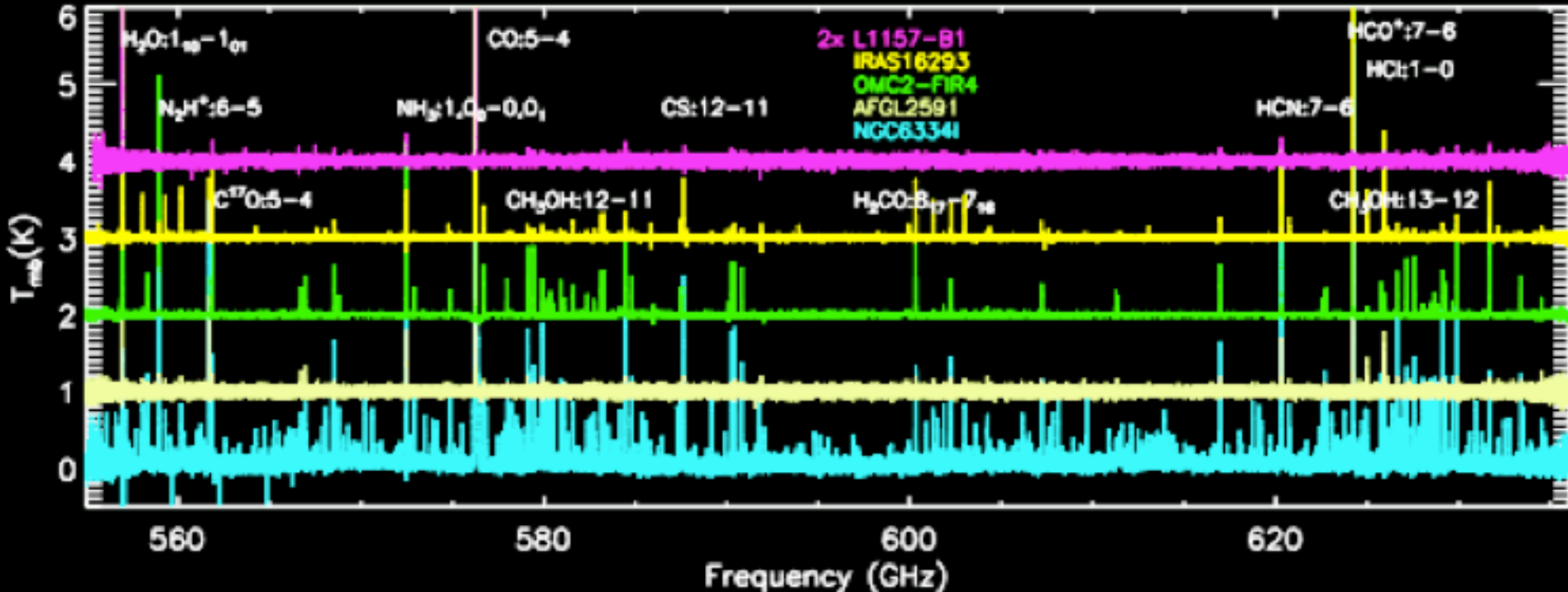
CHEMICAL HERSCHEL SURVEYS OF STAR FORMING REGIONS

News from CHESS

Ana López Sepulcre
(on behalf of C. Ceccarelli)

HDO/H₂O workshop. Garching, 15 Jan 2013

CHES overview: Band 1b



Ceccarelli, C., Bacmann, A., Boogert, A. et al. 2010, A&A, 521, 22

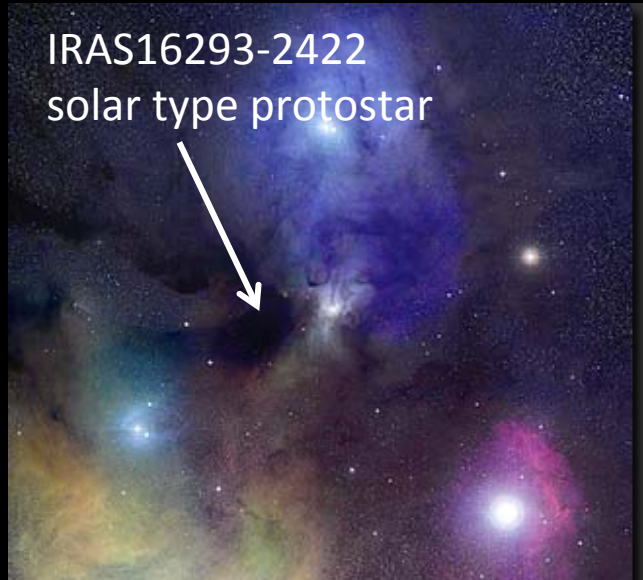
Herschel-HIFI unbiased spectral surveys of a sample of star forming regions covering a range of masses and evolutionary phases

CHES & HDO

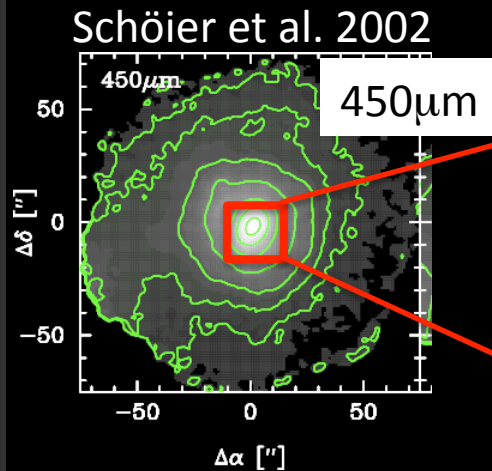
Source	Distance (pc)	Luminosity (L_{\odot})	Type
L1544	120	-	“Cold” Pre-Stellar Core
I16293E	120	-	“Warm” Pre-Stellar Core
L1157-B1	220	-	Outflow shock spot
IRAS16293-2422	120	21	Class0 low mass protostar
NGC1333-I4	250	40	Class0 low mass protostar
OMC2-FIR4	440	1×10^3	Intermediate mass protostar
AFGL2591	1000	2×10^4	High mass protostar
NGC6334I	1700	2×10^5	High mass hot core
NGC6334N	1700	2×10^4	High mass hot core
W51e	7000	2×10^6	High mass hot core

Studies of water deuteration only for 2 sources so far

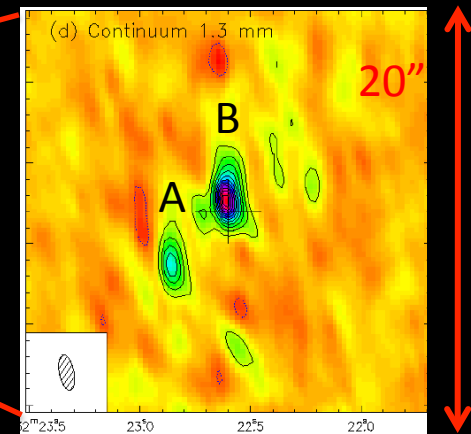
The low-mass protostar IRAS 16293-2422



$d = 120 \text{ pc}$



Bottinelli et al. 2004



2- M_{sun} ENVELOPE + BINARY SYSTEM +
a little of OUTFLOW in the **HIFI BEAM** ($\sim 36''$ @ 550GHz)

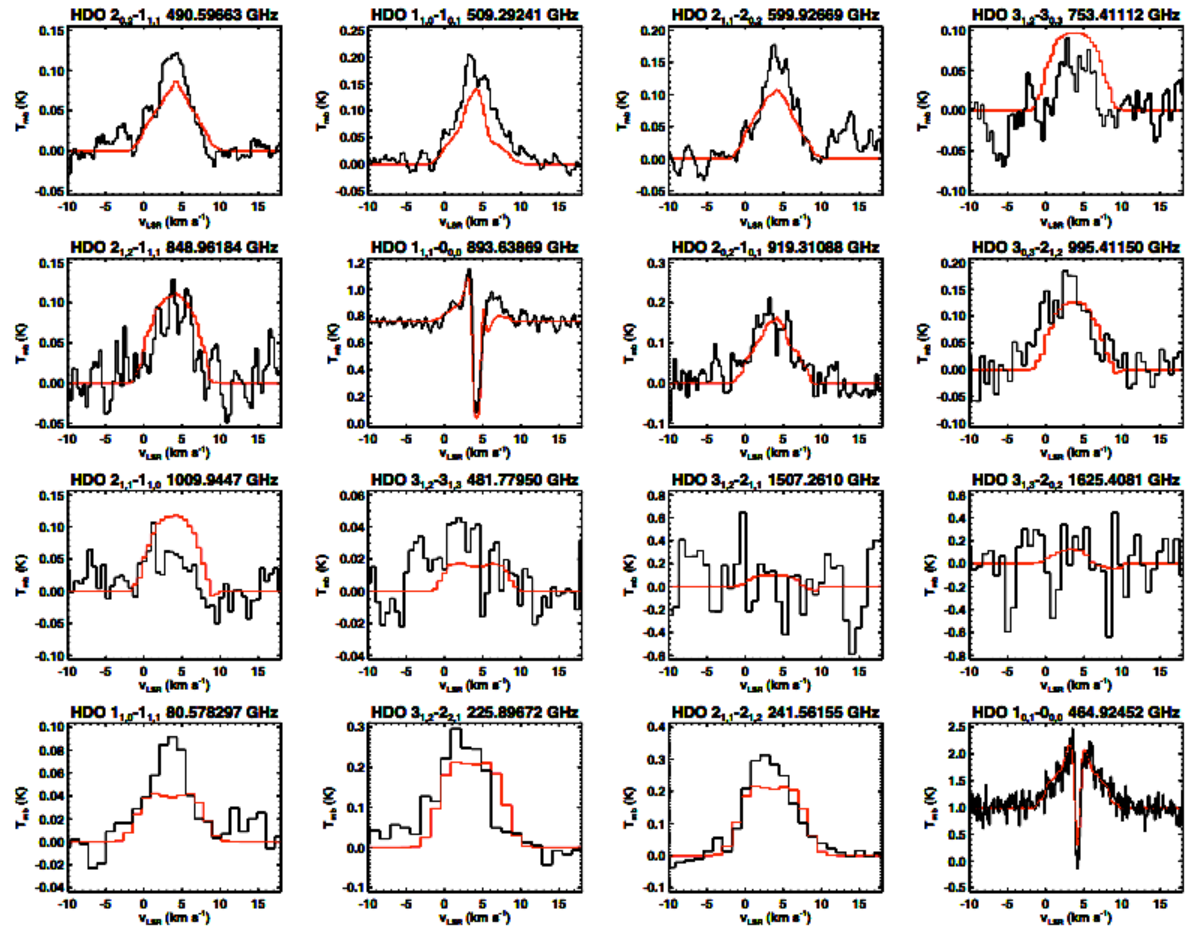
Dataset:

- CHES HIFI spectrum (480-1900 GHz)
- IRAM+JCMT spectral surveys: TIMASSS (3-0.8 mm)

The low-mass protostar IRAS 16293-2422

16 HDO lines
 $E_{\text{up}} = 22 - 170 \text{ K}$

15 H_2^{18}O lines
 $E_{\text{up}} = 53 - 450 \text{ K}$



Coutens et al. 2012,
A&A, 539, 132

OBSERVED and **MODELED** HDO SPECTRA

The low-mass protostar IRAS 16293-2422

Modelling: **RATRAN** code (Hogerheijde & van der Tak 2000)

Source structure from
Crimier et al. (2010)

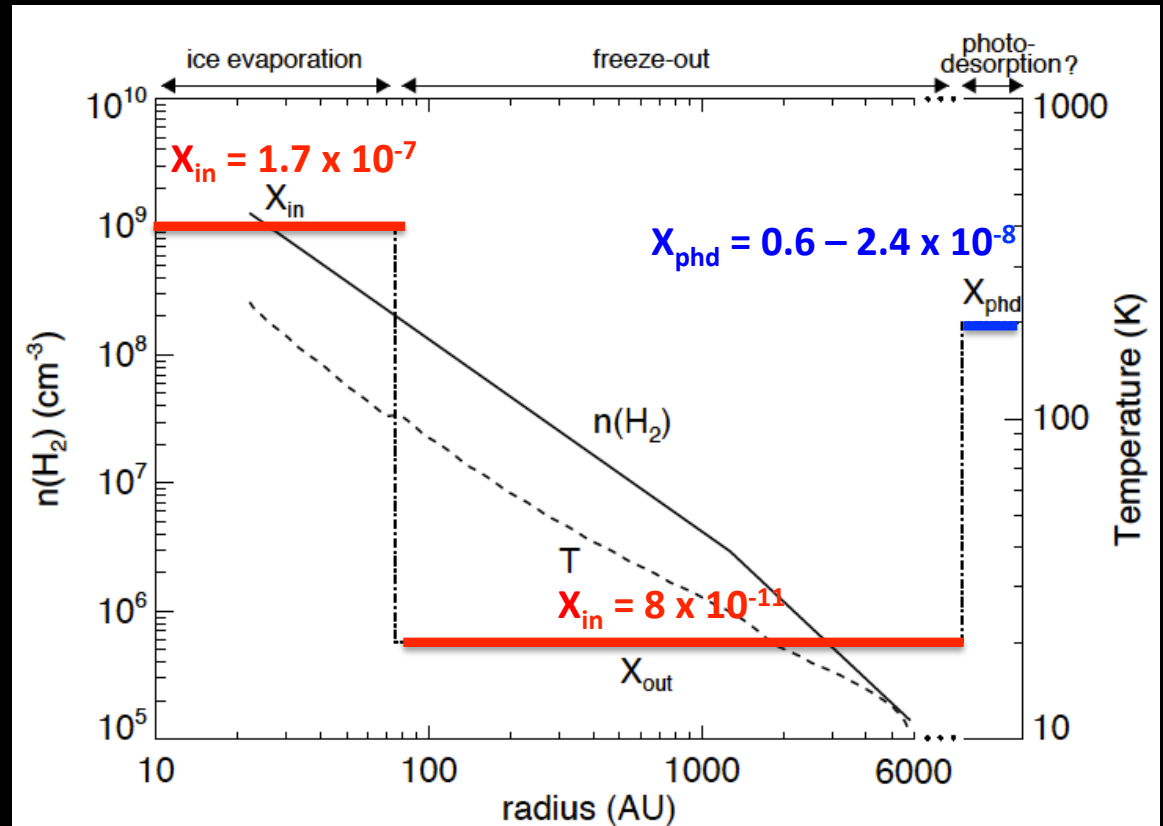
HDO/H₂O ratio:

Inner: 3.4 %

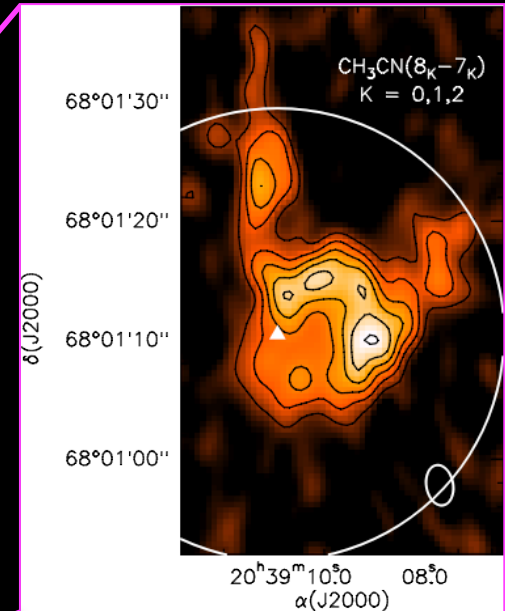
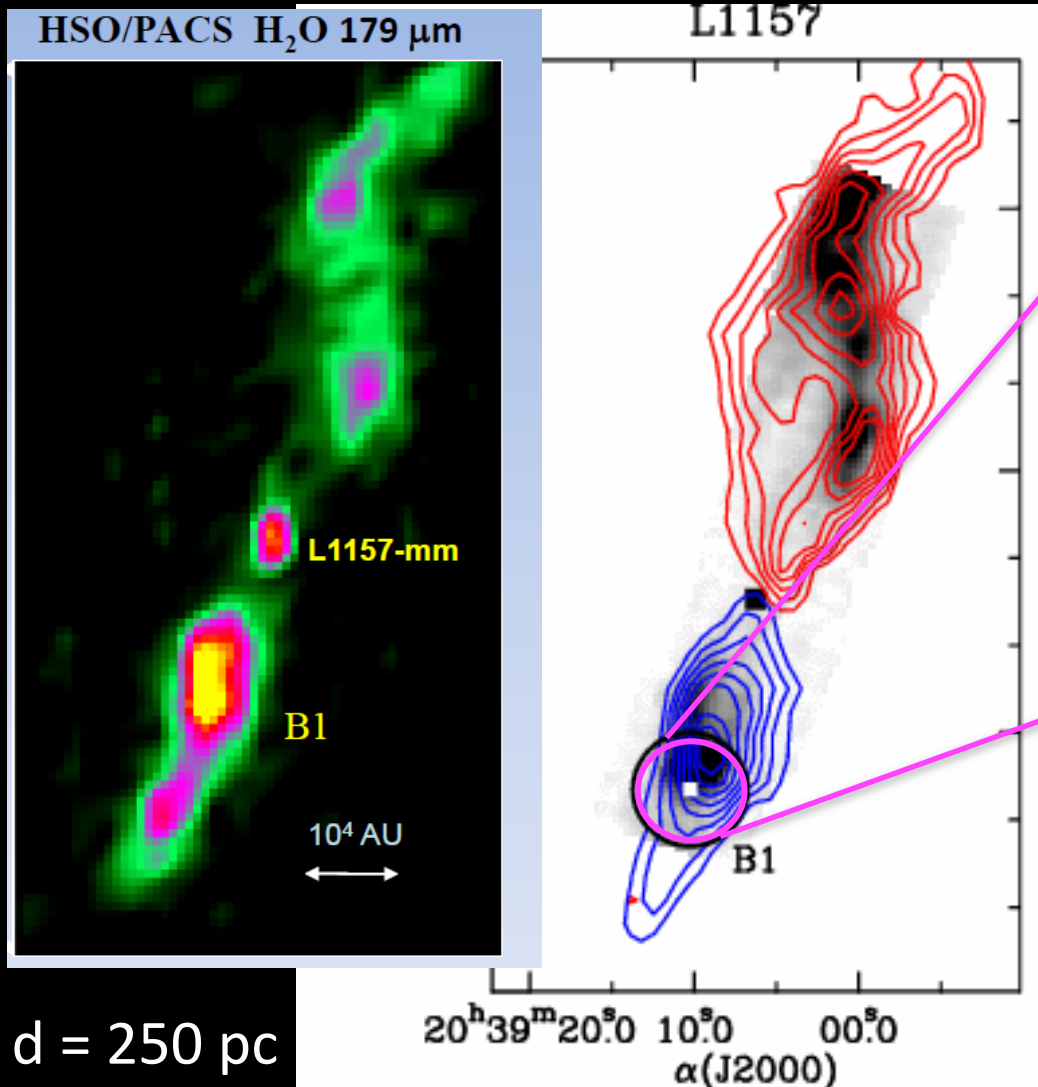
Outer: 0.5 %

Abs layer: ~ 4.8 %

Coutens et al. 2012,
A&A, 539, 132



The outflow shock spot L1157-B1

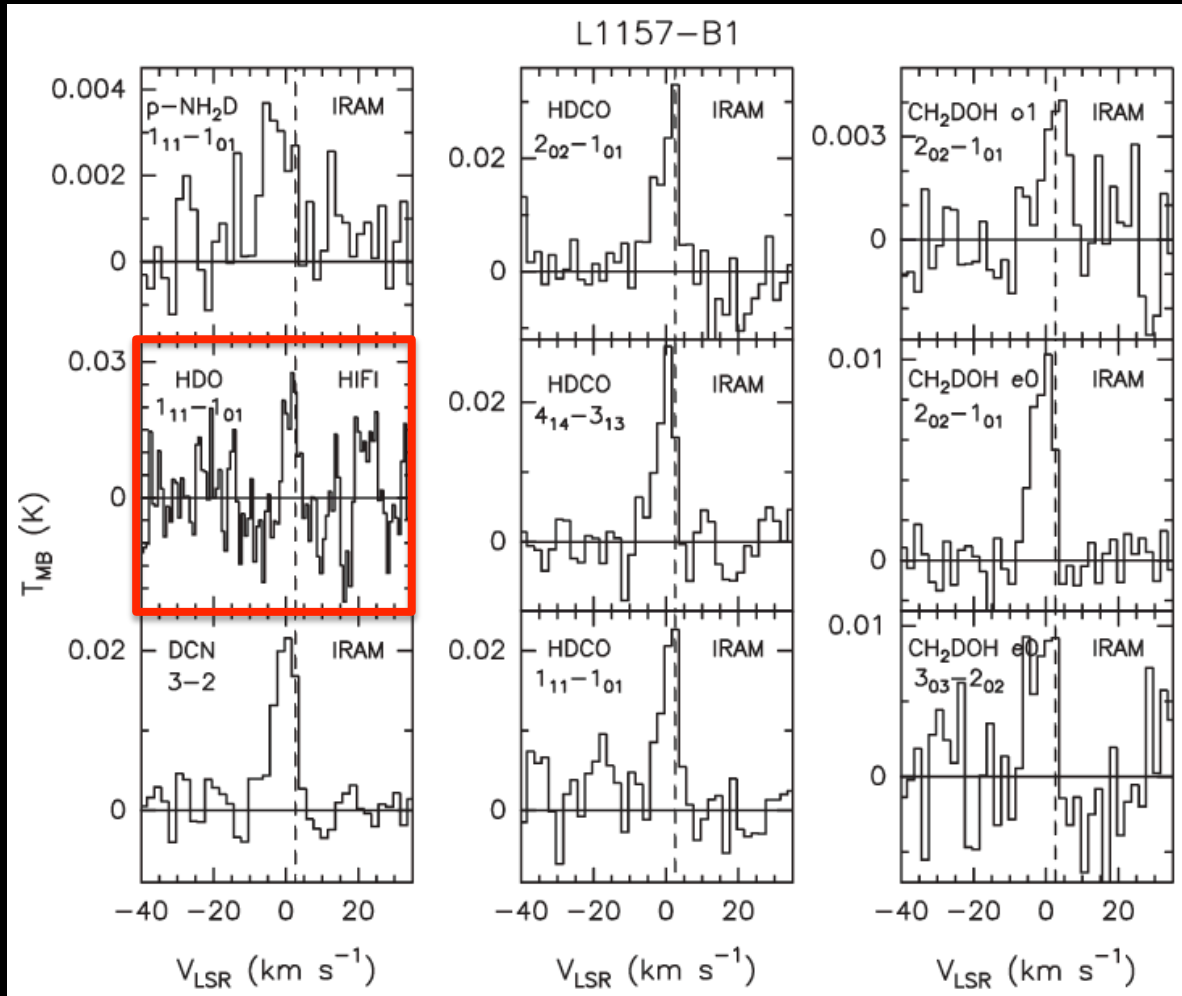


PdBI (Codella et al. 2009)

Dataset:

- CHES HIFI spectrum (480-1900 GHz)
- IRAM spectral surveys (3-0.8 mm)

The outflow shock spot L1157-B1



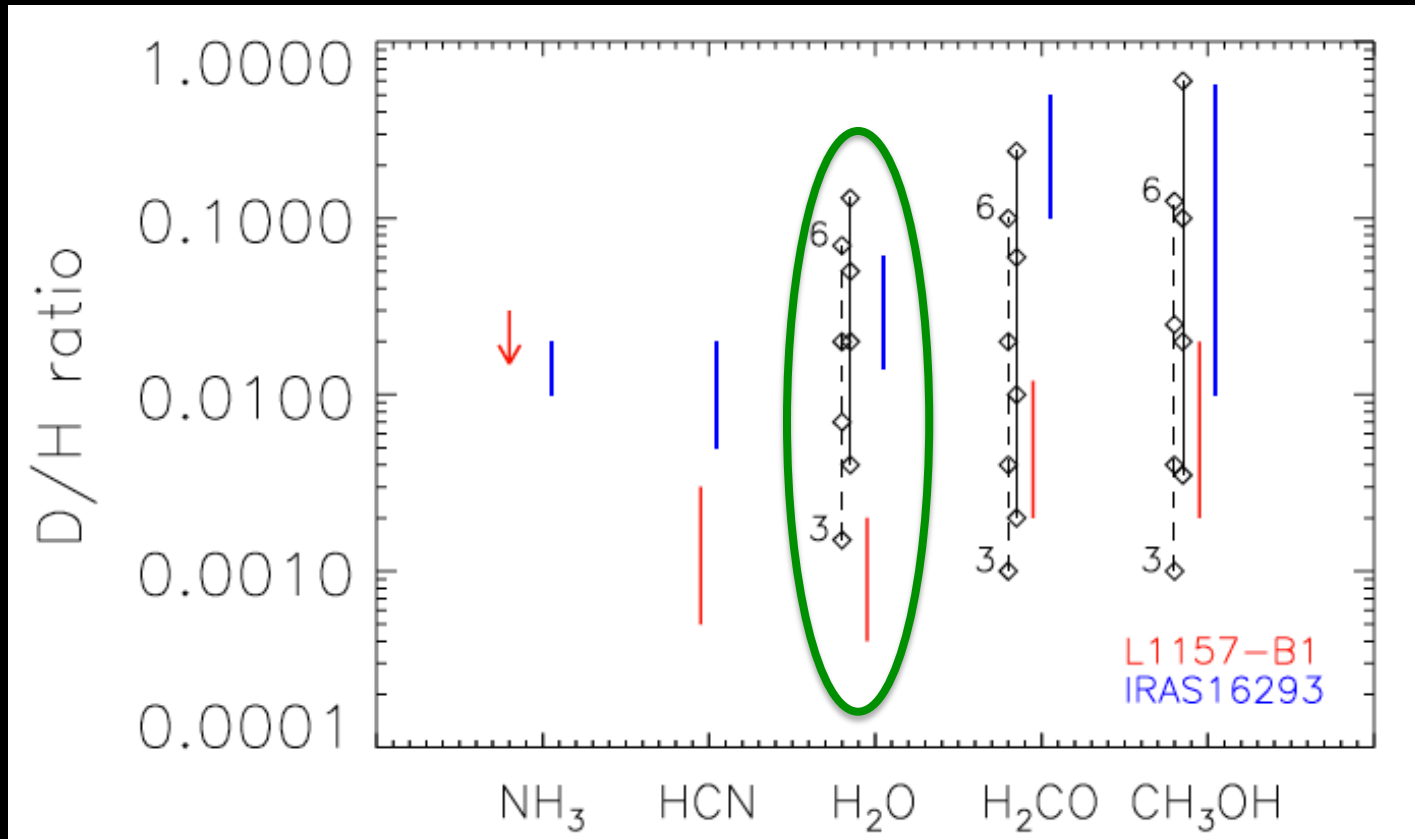
HIFI:

HDO(1₁₁-1₀₁) line
 p -H₂O(2₀₂-1₁₁) line

Column densities
derived using **LVG**
code (Ceccarelli et al.
2003)

HDO/H₂O ratio:
0.04 – 0.2 %

Comparison between the two sources



Codella et al. 2012, ApJL, 757, L9

IRAS 16293-2422: 0.5 - 5 %

L1157 - B1: 0.04 - 0.2 %

THANK YOU

